Added mass of an oscillating hemisphere at very-low and very-high frequencies

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Abstract

A floating hemisphere under forced harmonic oscillation at very-low and very-high frequencies is considered. The problem is reduced to an elliptic one, that is, the Laplace operator in the exterior domain with Dirichlet and Neumann boundary conditions. Asymptotic values of the added mass are found with an analytic prolongation for the surge mode, and with a semi-numerical computation with spherical harmonics for the heave mode. The general procedure is based on the use of spherical harmonics and its derivation is based on a physical insight rather than a mathematical one. This case can be used to test the accuracy achieved by numerical codes based on other formulations as finite or boundary elements.

1 Introduction

This work is concerned with a computation of the added mass of a floating hemisphere due to a forced oscillatory motion of the body in the free surface of an inviscid incompressible fluid. Two canonical cases are considered, namely those of heave and surge motions, where these nautical terms are used to describe a vertical or horizontal oscillation of the body.

The present problem has several applications. The oscillation of a floating body implies wave radiation so that it is of interest to wave energy conversion (e.g. see Falnes¹⁹). In seakeeping

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hydrodynamics (e.g. see Ohkusu,²⁴ Huang/Sclavounos³⁶), the hydrodynamic characteristics of a vessel due to action of surface waves when the wave motion is small enough compared to the body length can be computed with a linearized theory. Then, the response of a body to incident waves can be determined from the added-mass and damping coefficients associated to the forced motions of the body in the absence of waves, such as the heave motions of a semisubmersible platform due to the influence of ocean waves (see Hulme²). Besides, added mass effects can be recognized in ship vibrations. Although Lamb¹³ investigated the accelerated motion of a submerged cylinder, its relevance in ship vibrations was only properly recognized from the experimental work of Nicholls¹² and the mathematical one (using conformal mapping) of Lewis¹⁰.

From a physical point of view, the added mass effect comes from the inertia of the fluid. For instance, when a floating ship-like body performs a heave motion and the fluid is assumed as incompressible, there will be fluid motion between the hull and the free surface on the downbeat and back again on the upbeat, see Fig. 1. The added mass effects in this case come from the pressures transmitted to the hull arising from the inertia of the fluid. A similar analysis can be made for the surge motion, e.g. see Llloyd⁶ or Jennings³.

The formulation of problems involving floating hemispheres is analogous to that for the corresponding two-dimensional ones involving circular cylinders and perhaps, as Hulme¹ says, this is the reason why they have received rather comparatively little attention in the literature. The modern history of this subject began with Ursell⁸, who formulated and solved the boundary value problem for a semi-immersed heaving circular cylinder, where the velocity potential is represented as the sum of an infinite set of multipoles, each satisfying linearized free-surface boundary condition and each being multiplied by a coefficient determined by requiring the series to satisfy the kinematic boundary condition at a finite number of points on the cylinder. Grim²⁸ used a variation of the Ursell method to solve the two-parameter Lewis-form cylinders by conformal mapping onto a circle. Tasai⁷ and Porter³⁵ solved the added mass and damping using the Ursell approach for oscillating contours mappable onto a circle by the more general Theodorsen transformation, whereas Ogilvie³¹ computed the hydrodynamics forces on a completely submerged heaving circular cylinder. Frank³⁴ used an integral formulation where the velocity potential is represented by a distribution of sources over the submerged cross section; the density of the sources is an unknown function of the position along the contour to be determined from integral equations found by applying the kinematic boundary condition on the submerged part of the cylinder.

Havelock³³ gives an account of the heaving-hemisphere case, where the solution is found with a method similar to that used by Ursell for the circular cylinder, that is, the velocity potential is expressed in spherical polar coordinates as the sum of a wave source at the sphere center together with an infinite series of wave-free potentials. Then, the velocity potential satisfies all the boundary conditions except for the body surface, and the latter is used to generate an infinite linear system of equations for the infinite number of unknowns appearing in the expansion of the velocity potential. Hulme¹ used an essentially equivalent approach to Havelock's but with several modifications and more rigorous justification, where the solution is found by means of an expansion for the velocity potential in terms of an infinity series of spherical harmonics from which the relevant forces may be computed.

As it is usual in seakeeping flow problems, the forces exerted on oscillating bodies are given as the added mass and damping coefficients, which measure the components of the wave force in phase with the acceleration and velocity of the body respectively. In this work only the former case is considered. The proposed method is almost exact in the sense that the solution can be done with very high precision, and the results can be used to determine the accuracy achieved by other methods, e.g. finite or boundary elements, which can be used to treat more general body geometries, e.g. see Nigro²⁵ et al.. Storti^{26,27} et al., D'Elía^{17,18,16} et al..

In this work, the heave and surge modes at very low and very high frequencies of the unit hemisphere are reformulated as boundary value problems extended to all the space, they are solved by orthogonal expansion by means of spherical harmonics. In Hulme's work, the derivation of the asymptotic form for the surge-mass coefficient at very high frequencies ($\omega \to \infty$) is, as Hulme says, *suggestive* rather than conclusive, so an alternative derivation is given here for the same coefficient where the present estimation is closely related to Hulme's estimate.

2 The oscillating unity hemisphere

An oscillating unity hemisphere in a forced motion is considered. The unit hemisphere is the open surface $r = 1, 0 \le \theta \le \pi/2, 0 \le \varphi \le 2\pi$, where r is the radius, θ is the azimuthal angle and φ is the circumferential one. Its edge is on the free surface of an irrotational and incompressible fluid without a mean flow, the fluid depth is assumed as infinity, the z-axis positive downward and the hydrostatic equilibrium plane is z = 0, see Figs. 2 and 3. Due to the symmetry, the spherical coordinates (r, θ, φ) are employed. In this work a whole linearized analysis is performed and for this reason the linearized surface boundary condition at z = 0 is only used. As it is well known (e.g. see Newman²¹, Ogilvie³², Ohkusu²⁴, Stoker²⁰), the standard free surface boundary condition of linearized water-wave theory for sinusoidal time dependence is given by

$$\phi_{,n} = \frac{\omega^2}{g}\phi \quad \text{at } z = 0; \tag{1}$$

where ϕ is the velocity potential, $\phi_{,n} \equiv \partial \phi / \partial n$ is the normal velocity, n is the unit normal, z = 0is the hydrostatic free surface, ω is the circular frequency of the sinusoidal oscillation, and g is the gravity acceleration.

2.1 Limits at very-low and very-high frequencies

In both limit processes $\omega \to 0$ and $\omega \to \infty$ for the linearized free-surface boundary condition given by Eq. 1, it will be assumed that the velocity potential ϕ , the normal velocity $\phi_{,n}$ and the gravity acceleration g remain finite and bounded. Then, on the one hand, at very-low frequencies $(\omega \to 0)$ the free-surface boundary Eq. 1 shrinks to the homogeneous Neumann one $\phi_{,n} = 0$ at z = 0, that is,

$$\omega \to 0 : \begin{cases} \Delta \phi = 0 & \text{in } \Omega; \\ \phi_{,n} = 0 & \text{at } z = 0; \\ \phi_{,n} = h & \text{at } \Gamma_b; \\ |\phi| \to 0 & \text{for } |\boldsymbol{x}| \to \infty. \end{cases}$$
(2)

On the other hand, at very-high frequencies $(\omega \to \infty)$ the free-surface boundary Eq. 1 shrinks to the homogeneous Dirichlet one $\phi = 0$ at z = 0, then

$$\omega \to \infty : \begin{cases} \Delta \phi = 0 & \text{in } \Omega; \\ \phi = 0 & \text{at } z = 0; \\ \phi_{,n} = h & \text{at } \Gamma_b; \\ |\phi| \to 0 & \text{for } |\boldsymbol{x}| \to \infty. \end{cases}$$
(3)

In Eq. 2-3, Δ is the Laplace operator, $h = h(\theta, \varphi)$ is the load given by the normal displacement of the mode under consideration, and the last condition is the radiation boundary condition at infinity. For simplicity, it is assumed that the load $h = h(\theta, \phi)$ is real, that is, the body motion is in phase with the fluid velocity. It should be noted that the original linearized flow problem is defined only in the lower region $z \leq 0$ but the analysis performed for the two limit processes of the linearized surface boundary condition at z = 0, that is, for $\omega \to 0$ [Very Low Frequencies (VLF)] and for $\omega \to \infty$ [Very High Frequencies (VHF)], suggests that the boundary conditions $\phi_{,n} = 0$ at VLF and $\phi = 0$ at VHF respectively could be taken into account at the plane z = 0.

2.2 The heave and surge modes

The heave-mode excitation of the body (vertical oscillation) produces a symmetrical displacement around the vertical axis z and it can be written as $h = \cos \theta$, with $0 \le \theta \le \pi/2$, see Fig. 4. Analogously, the surge-mode excitation (horizontal oscillation) produces an antisymmetrical displacement with respect to the plane x = 0 and it can be written as $h = \sin \varphi$, with $0 \le \varphi < 2\pi$, see Fig. 5. Once the velocity potential ϕ is solved for each mode (i.e. the surge ϕ^1 and the heave ϕ^3 ones), the added mass A_{ij} in the *i* degree of freedom due to a harmonic unity excitation on the *j*-direction is computed as the surface integral

$$A_{ij} = -\rho \int_{\Gamma_b} d\Gamma \phi^i \phi^j_{,n} ; \qquad (4)$$

over the body surface Γ_b , where i = j = 1 for the surge mode and i = j = 3 for the heave one, e.g. see Ohkusu,²⁴ Newman.²¹ Then, the added mass coefficient is obtained as $A'_{ij} = A_{ij}/(\rho V)$, where $V = (2/3)\pi R^3$ is the body volume of the hemisphere and ρ is the fluid density.

3 The extended flow problems

By symmetry, Eqn. 2-3 can be reproduced extending the flow problem to upper region z < 0by means of a reflection with respect to the plane z = 0 and extending the load $h = h(\theta, \phi)$ in an appropriate way. For instance, the homogeneous Neumann boundary condition is obtained when the load h is extended in a symmetrical way, i.e. h(x, y, z) = h(x, y, -z), while the homogeneous Dirichlet one is obtained when the load h is extended in an anti-symmetrical way, i.e. h(x, y, z) = -h(x, y, -z). It should be noted that the plane z = 0 is not really necessary in the subsequent analysis since the surface load extension *automatically* satisfied the suggested linearized boundary conditions $\phi_{,n} = 0$ at VLF and $\phi = 0$ at VHF, respectively.

3.1 The extended heave-modes at VLF and VHF

The extended heave loads at Very Low Frequencies (VLF) and Very High Frequencies (VHF) are obtained from the sphere ones as

$$\begin{cases} h = |\cos \theta| & \text{for low frequencies } (\omega \to 0); \\ h = \cos \theta & \text{for high frequencies } (\omega \to \infty); \end{cases}$$
(5)

where now $0 \le \theta \le \pi$ due to the extension to upper region z < 0. Then, the boundary value problem of the heave-mode at very-low frequencies ($\omega \to 0$) is written as

extended heave-mode at VLF
$$(\omega \to 0)$$
 :

$$\begin{cases}
\Delta \phi = 0 & \text{in } \Omega'; \\
\phi_{,n} = |\cos \theta| & \text{at } \Gamma_b'; \\
\phi_{,n} = 0 & \text{at } z = 0; \\
|\phi| \to 0 & \text{for } |\mathbf{x}| \to \infty;
\end{cases}$$
(6)

see Fig. 6, where $\Omega' = \Omega_e \cup \Omega'_e$ is the extended flow domain, Ω_e and Ω'_e are the flow domain exterior to the hemisphere and its extension through the reflection plane z = 0, respectively, and Γ'_b is extended hemisphere surface through the same plane. In general, due to the module on source term $|\cos \theta|$, this case does not have a closed solution. Therefore it must be found with other resources like spherical harmonics, as considered in this work. Analogously, the boundary value problem of the heave-mode at very-high frequencies ($\omega \to \infty$) is written as

extended heave-mode at VHF
$$(\omega \to \infty)$$
 :
$$\begin{cases} \Delta \phi = 0 & \text{in } \Omega'; \\ \phi_{,n} = \cos \theta & \text{at } \Gamma'_b; \\ \phi = 0 & \text{at } z = 0; \\ |\phi| \to 0 & \text{for } |\mathbf{x}| \to \infty; \end{cases}$$
(7)

see Fig. 7 where, since the free surface boundary condition for VHF is $\phi = 0$, its right hand side term has been extended in an anti-symmetric way and, then, it is equivalent to a sphere in infinite medium.

3.2 The extended surge-mode

On the other hand, the extended surge loads at the VLF and VHF, e.g. see Fig. 8 and 9, are

$$\begin{cases} h = \sin \varphi \sin \theta & \text{for low frequencies } (\omega \to 0); \\ h = \sin \varphi \sin \theta \, \text{sign} \{\cos \theta\} & \text{for high frequencies } (\omega \to \infty). \end{cases}$$
(8)

Then, the boundary value problem of the surge-mode at very-low frequencies $(\omega \to 0)$ is written as

extended surge-mode at VLF
$$(\omega \to 0)$$
 :

$$\begin{cases}
\Delta \phi = 0 & \text{in } \Omega'; \\
\phi_{,n} = \sin \varphi \sin \theta & \text{at } \Gamma'_{e}; \\
\phi_{,n} = 0 & \text{at } z = 0; \\
|\phi| \to 0 & \text{for } |\mathbf{x}| \to \infty;
\end{cases}$$
(9)

and the boundary value problem at very-high frequencies $(\omega \to \infty)$ is written as

extended surge-mode at VHF
$$(\omega \to \infty)$$
 :
$$\begin{cases} \Delta \phi = 0 & \text{in } \Omega'; \\ \phi_{,n} = \sin \varphi \sin \theta \, \text{sign} \{\cos \theta\} & \text{at } \Gamma'_e; \\ \phi = 0 & \text{at } z = 0; \\ |\phi| \to 0 & \text{for } |\mathbf{x}| \to \infty. \end{cases}$$
(10)

4 Solution of the extended flow problems

The extended flow problems can be solved in an analytical way or by series. The solutions for the extended flow problems corresponding to the heave-mode at VHF ($\omega \to \infty$) and surge-mode at VLF ($\omega \to 0$), are the same of a sphere in an infinity medium and uniform velocity, so the additional mass is half the displaced mass, that is,

$$\begin{cases}
A_{33}(\omega \to \infty) = \frac{\pi}{3}\rho R^3; \\
A_{11}(\omega \to 0) = \frac{\pi}{3}\rho R^3.
\end{cases}$$
(11)

In the other two cases, the heave-mode at VLF ($\omega \to 0$) and the surge-mode at VHF ($\omega \to \infty$), are not that easy to obtain, so the solutions are found expanding the sources by means of spherical harmonics.

5 Spherical harmonics

The exterior potential problem

$$\begin{cases} \Delta \phi = 0 & \text{for } r > 1; \\ \phi = f(\theta, \varphi) & \text{at } r = 1; \end{cases}$$
(12)

where $\phi = \phi(\theta, \varphi)$ is solved expanding the function $f(\theta, \varphi)$ in terms of the harmonics

$$f(\theta,\varphi) = \sum_{n=0}^{\infty} a_{n0} P_n(\cos\theta) + \sum_{n=0}^{\infty} \sum_{m=1}^{n} [a_{nm} \cos(m\varphi) + b_{nm} \sin(m\varphi)] P_n^m(\cos\theta) ;$$
(13)

where $P_n(u)$ are the Legendre polynomials, with $u = \cos \theta$, $P_n^m(u) = (1 - u^2)^{m/2} d^m P_n/du^m$ are the associated ones, while the coefficients are given by

$$a_{n0} = \frac{2n+1}{4\pi} \int_{r=1} d\Gamma f(\theta,\varphi) P_n(\cos\theta) ; \qquad (14)$$

$$a_{nm} = \frac{2n+1}{2\pi} \frac{(n-m)!}{(n+m)!} \int_{r=1} d\Gamma f(\theta,\varphi) P_n^m(\cos\theta) \cos m\varphi ; \qquad (15)$$

$$b_{nm} = \frac{2n+1}{2\pi} \frac{(n-m)!}{(n+m)!} \int_{r=1} d\Gamma f(\theta,\varphi) P_n^m(\cos\theta) \sin m\varphi ; \qquad (16)$$

where $d\Gamma = \sin \theta \ d\varphi \ d\theta$ is the differential of the solid angle in spherical coordinates. Once this expansion is computed, the exterior potential can be written as

$$\phi(r,\theta,\varphi) = \sum_{n=0}^{\infty} Y_n(\theta,\varphi) r^{-(n+1)}; \qquad (17)$$

where

$$Y_n(\theta,\varphi) = a_{n0}P_n(\cos\theta) + \sum_{m=1}^n \left[a_{nm}\cos(m\varphi) + b_{nm}\sin(m\varphi) \right] P_n^m(\cos\theta) .$$
(18)

Then, the Neumann problem can be solved taking derivatives with respect to r and evaluating at r = 1, that is, $h \equiv \phi_{,r}|_{r=1}$, obtaining the expression

$$h(\theta,\varphi) = -\sum_{n=0}^{\infty} (n+1) Y_n(\theta,\varphi) ; \qquad (19)$$

from which analogous relations are obtained

$$a_{n0} = \frac{2n+1}{4\pi(n+1)} \int_{r=1} d\Gamma \ h(\theta,\varphi) \ P_n(\cos\theta) \ ; \tag{20}$$

$$a_{nm} = \frac{2n+1}{2\pi(n+1)} \frac{(n-m)!}{(n+m)!} \int_{r=1} d\Gamma \ h(\theta,\varphi) \ P_n^m(\cos\theta) \ \cos(m\varphi) \ ; \tag{21}$$

$$b_{nm} = \frac{2n+1}{2\pi(n+1)} \frac{(n-m)!}{(n+m)!} \int_{r=1} d\Gamma \ h(\theta,\varphi) \ P_n^m(\cos\theta) \ \sin(m\varphi) \ . \tag{22}$$

Once the coefficients of the expansion are obtained, the added mass can be found from

$$A_{jj} = -\rho \int_{r=1}^{\infty} d\Gamma \phi \phi_{,r} = -\rho \int_{0}^{2\pi} G(\varphi) d\varphi \int_{0}^{\pi} H(\theta) d\theta ; \qquad (23)$$

where $G(\varphi) = \{1, \sin^2 \varphi\}$ for the heave and surge motions, respectively, and

$$H(\theta) = \sum_{n=0}^{\infty} \left\{ 2a_{n0}^2 \left[P_n(\cos\theta) \right]^2 + \sum_{m=1}^n \left(a_{nm}^2 + b_{nm}^2 \right) \times \left[P_n^m(\cos\theta) \right]^2 \right\} ;$$
(24)

where the orthogonality property of the spherical harmonics was taken into account. Finally, using properties of the Legendre polynomials

$$A_{jj} = \rho \alpha \sum_{n=1}^{\infty} \frac{n+1}{2n+1} \left[2a_{n0}^2 + \sum_{m=1}^n (a_{nm}^2 + b_{nm}^2) \frac{(n+m)!}{(n-m)!} \right] ; \qquad (25)$$

where $\alpha = \{2\pi, \pi\}$, for the heave and surge motions, respectively.

6 Hemisphere in heave-mode at very-low frequencies

The load in the heave-mode at VLF is $h(\theta, \varphi) = |\cos \theta|$, so

$$a_{n0} = \frac{2n+1}{4\pi(n+1)} \int_{-1}^{1} |\mu| P_n(\mu) d\mu \int_{0}^{2\pi} d\varphi ; \qquad (26)$$

where $\mu = \cos \theta$. As the $P_n(\mu)$ are even (odd) for *n* even (odd), only remains its even terms and then

$$a_{n0} = \frac{2n+1}{n+1} \int_0^1 \mu P_n(\mu) d\mu ; \qquad (27)$$

for *n* even. For computing this integral, the P_n terms are generated in a recursive way with the initial conditions $P_0 = 1$, $P_1 = \mu$, and the next terms P_2 , ..., P_n are obtained by solving

$$(n+1) P_{n+1} - (2n+1) \mu P_n(\mu) + n P_{n-1}(\mu) = 0.$$
(28)

The coefficients of the polynomials $\mu P_n(\mu)$ are obtained from the $P_n(\mu)$ ones, and the integral is made in a semi-analytical way. The final result for the added mass in the heave-mode at VLF for the sphere is

$$A_{33}(\omega \to 0) = 1.740 \ 335 \ 785 \ 143 \ \rho \ R^3 \ ; \tag{29}$$

corresponding to $A'_{33}(\omega \to 0) = 0.830$ 949 128 536, that is, the non-dimensional coefficient with respect to the hemisphere mass $2/3\pi\rho R^3$.

7 Hemisphere in surge-mode at very-high frequencies

In this case, due to the load $h = \sin \varphi \sin \theta \text{sign} \{\cos \theta\}$, the only no-null coefficients are the b_{n1} terms. For obtaining them an integral from $\mu = -1$ to $\mu = 1$ must be made, with a function which includes the $P_n^1(\mu)$ terms. These terms have a factor $\sqrt{1-\mu^2}$, so it is more convenient to perform a semi-numerical integration (e.g. see Appendix). The final result is

$$A_{11}(\omega \to \infty) = 0.570 \ 136 \ 261 \ 149 \ \rho R^3 ; \tag{30}$$

corresponding to $A'_{11}(\omega \to \infty) = 0.272\ 220\ 012\ 593$, that is, the non-dimensional coefficient with respect to the hemisphere mass $(2/3)\pi\rho R^3$.

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8 Discussion

As it can be seen, the solutions to these problems have somewhat different properties according to whether the body oscillates in heave (vertically) or in surge (horizontally). On the one hand, at Very Low Frequencies (VLF), if the body oscillates in heave its image moves oppositely (see Fig. 6) so that the two act together somewhat as a *pulsating source*, while if the body oscillates in surge its image moves in the same direction (see Fig. 8) and the total effect on the pressure load is the same as for a *horizontal dipole* in an infinite fluid. On the other hand, at Very High Frequencies (VHF), if the body oscillates in heave its image moves in the same direction (see Fig. 7) so that the two act together somewhat as a *vertical dipole*, while if the body oscillates in surge its image moves oppositely (see Fig. 9) and the total effect on the pressure load is the same as a *pair* of *horizontal dipoles* oriented in *opposite* directions. Furthermore, the intensity of the velocity potentials ϕ^j , with j = 1 (surge) and j = 3 (heave), are proportional to the source terms of the corresponding governing differential equations which, in turn, are fixed by these pressures loads. As the added mass is proportional to the integral of the fluid velocity at the surface of the extended body, then the resulting added mass can be expected to be greater for heave motion at Very Low Frequencies (VLF) (since its pressure load is always positive on the extended surface) than for the surge one at Very High Frequencies (VHF) (since its pressure load has both positive and negative values). In any case, the added mass at any frequency in heave motion is always greater than in the surge one (since the displaced fluid is greater in heave than in surge), as it is well known from the corresponding plots as functions of the frequency obtained, for instance, by numerical computations, e.g. see Papanikolau⁴ or D'Elía¹⁵. A comprehensive analysis of relations between added masses and sources and doublets are given, for example, by $Ogilvie^{32}$ and Landweber²².

The present estimates for the added mass coefficients with respect to the hemisphere mass $2/3\pi\rho R^3$, for the surge mode (i = 1, longitudinal oscillation) and for the heave one (i = 3, vertical oscillation), at VLF ($\omega \rightarrow 0$) and VHF ($\omega \rightarrow \infty$) limits, are summarized in Table 1 and compared to some literature values found (i) for the surge/sway mode, e.g. see Sierevogel²³, Prins¹⁴ (where only the intervals [0.25, 1.50] and [0.6, 1.5] are respectively considered and so the

extrapolations are rather doubtful); (ii) for the heave one, e.g. see Korsmeyer¹¹ and Liapis³⁰; and (iii) Hulme¹. The Sierevogel, Prins and Liapis results are obtained with a panel method and Kelvin source; Korsmeyer used a panel method with Fourier transform and complex impedance extended to very-low frequencies, while Hulme's numerical results are obtained by spherical harmonics but with a rather different derivation and implementation.

In Hulme's work, the solution for each mode is expressed in terms of infinite series of spherical harmonics and then it is found solving truncated infinite linear system of equations. Hulme notes that its procedure is successful for slow oscillations of the body, i.e. small Ka, where a is the hemi-sphere radius (here Hulme's notation is employed) but, for high frequencies Ka the system is ill-conditioned so, for such cases, the problems are re-formulated as integral equations whose kernels become small as $Ka \to \infty$. This same procedure is also used by Ursell⁹ and Davis⁵ but, as a rigorous treatment of the surge case involves a significant amount of mathematical labour, Hulme gives a plausible procedure from which the surge-added mass coefficient is approximated as

$$A^{(1)} \sim C_1 - \frac{C_2}{Ka} \qquad \text{as } Ka \to \infty;$$
(31)

with

$$C_1 = -\frac{3}{2} \sum_{n=1}^{\infty} \alpha_n I\{2n, 1; 1\} \quad ; \quad C_2 = +\frac{3}{2} \sum_{n=1}^{\infty} \beta_n I\{2n, 1; 1\} \; ; \tag{32}$$

(Eq. 5.14 to 5.16, p.460, op. cit.) where the terms in the series for C_1 and C_2 decay as $1/n^3$ and $\log(n)/n^2$, respectively, so the C_2 coefficient is neglected by Hulme and, then, the surge-mass coefficient at VHF is assumed equal to $A^{(1)} \approx C_1 \approx 0.273$ 239... which is closely related to the present estimate.

As conclusion, an alternative derivation of the added mass coefficients for the heave and surge motions of the unit floating hemisphere at very-low and very-high frequencies was shown. As a particular case, the numerical value obtained by the present procedure for the surge-mass coefficient at very high frequencies ($\omega \to \infty$) is closely related to Hulme's *suggestive* one. Other modeling efforts would be focused on its extension to other analytic body shapes which can be extended to the upper region as spheroids or ellipsoids. These simpler geometries can be used as test cases for validation of related numerical codes, as those used in ship-hydrodynamics or fluid-solid interaction.

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Appendix

The load in the surge-mode at VHF is $h = \sin \varphi \sin \theta \, \text{sign} \{\cos \theta\}$. The only no-null coefficients are the b_{n1} terms, that is,

$$b_{n1} = \frac{2n+1}{2\pi} \frac{1}{n(n+1)^2} \int_0^{2\pi} \sin^2 \varphi d\varphi \int_0^{\pi} \sin \theta \, \operatorname{sign}\{\cos \theta\} P_n^1(\cos \theta) \sin \theta d\theta \; ; \tag{33}$$

introducing $\mu = \cos \theta$ and integrating in φ ,

$$b_{n1} = \frac{2n+1}{2} \frac{1}{n(n+1)^2} \int_{-1}^{1} \sqrt{1-\mu^2} \operatorname{sign}\{\mu\} P_n^1(\mu) d\mu ; \qquad (34)$$

only the even terms are no-nulls, so

$$b_{n1} = \frac{2n+1}{n(n+1)^2} \int_0^1 (1-\mu^2) \frac{dP_n}{d\mu} d\mu ; \qquad (35)$$

integrating by parts,

$$b_{n1} = \frac{2n+1}{n(n+1)^2} \left[\left| (1-\mu^2) P_n(\mu) \right|_0^1 + 2 \int_0^1 \mu P_n(\mu) d\mu \right] ; \qquad (36)$$

or

$$b_{n1} = \frac{2n+1}{n(n+1)^2} \left[-P_n(0) + 2\int_0^1 \mu P_n(\mu) d\mu \right] ; \qquad (37)$$

which is computed with a semi-numerical procedure, for instance, with Octave²⁹.

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Figure 1: Sketch of the movement of an incompressible fluid in ship-like vibration due to a heave motion.



Figure 2: Geometrical description of a seakeeping-like flow: original problem (left) and extended to the upper plane (right).



Figure 3: Cartesian coordinates (x, y, z) and spherical ones (r, θ, φ) .



Figure 4: The heave load $h(\theta, \varphi) = \cos \theta$, with $0 \le \theta \le \pi/2$.



Figure 5: The surge load $h(\theta, \varphi) = \sin \varphi \sin \theta$, with $0 \le \varphi < 2\pi$ and $0 \le \theta < \pi/2$.



Figure 6: Symmetrical load extension $h = |\cos \theta|$ for the heave-mode at very-low frequencies $\omega \to 0$.



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Figure 8: The load extension $h = \sin \varphi \sin \theta$ for the surge-mode at Very Low Frequencies (VLF) $(\omega \to 0)$.



Figure 9: The load extension $h = \sin \varphi \sin \theta \, \text{sign} \{\cos \theta\}$ for the surge-mode at Very High Frequencies (VHF) $(\omega \to \infty)$.

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VLF $(\omega \to 0)$		VHF $(\omega \to \infty)$		
surge	heave	surge	heave	
A'_{11}	A'_{33}	A'_{11}	A'_{33}	reference
0.5	0.8	0.20	0.40	Sierevogel ²³ , Prins ¹⁴
0.5	0.8	0.25	0.45	Korsmeyer ¹¹ , Liapis ³⁰
0.5	$0.830\ 951$	$0.273\ 239$	0.50	Hulme^1
0.5	$0.830\ 930$	$0.272\ 220$	0.50	present work

Table 1: Added mass at Very-Low Frequencies (VLF) ($\omega \rightarrow 0$) and Very-High Frequencies (VHF) ($\omega \rightarrow \infty$) for the surge mode i = 1 (longitudinal oscillation) and the heave one i = 3 (vertical oscillation) on the unit hemisphere.