# THEORY OF AFFINE SHELLS: HIGHER ORDER ESTIMATES 

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#### Abstract

The classical Theory of Shells has been exposed, through the contributions of many authors, within the framework of Euclidean Geometry, i.e., based on the classical theory of surfaces in three-dimensional space, which is invariant under translations and rotations. For diverse viewpoints of presentation see (F. John, Comm. Pure Appl. Math. 18(1/2): 235-267 (1965); 24(5): 583-615 (1971)) and other references therein. More recently, we ourselves have been working in a new development of this theory based, from the geometrical viewpoint, in those objects which remain invariant under the action of the Unimodular Affine Group, i.e., dealing with Affine Surface Geometry (S. Gigena et al., Mec. Comp., 21: 18621881 (2002); 22: 1953-1963 (2003); 23: 639-652 (2004); 24: 2745-2758 (2005)). In this paper we study exclusively the behavior of physical objects of the shell in the interior, without reference to any boundary conditions at the edge. For the interior behavior one needs as the only tool a certain kind of a priori estimates. These interior regularity estimates, similar to those occurring in the theory of Partial Differential Equations, rigorously assign a definite order of magnitude to every quantity occurring in the theory. Our main goal here is to establish those estimates for the strain and stress tensors, as well as for the higher order covariant derivatives of both, within the framework of the Theory of Affine Shells.


## 1 INTRODUCTION

The Theory of Shells is a scientific and technological topic with quite a rich history and many, diverse applications to the real world: Engineering, Industry, Avionics, and so on. The usual viewpoint of presentation, which is exposed indeed in the great majority of texts and research articles, makes use of classical, Euclidean geometry of surfaces in three-dimensional space, particularly with regards to the invariants of the Euclidean group, ASO ( $3, \mathbb{R}$ ), i.e., the group of transformations generated by translations and rotations of the space. See, for instance, John (1965 and 1971), Koiter (1970), Love (1944), Möllmann (1981). Within that context what it is called "normal" is the Euclidean one, and the "distance" is the measure with respect to the norm induced by the usual scalar product of vectors (positive definite), which is the main, fundamental invariant in Euclidean geometry, as exposed, for example, in the book by Millman and Parker (1977).

On our part, for the latter few years we have been working on an alternative foundation, exposure and development of the theory of shells which is invariant, from the geometrical point of view, under the action of the unimodular affine group, ASL ( $3, \mathbb{R}$ ). See Gigena et al. (2002, 2003, 2004, and 2005) for full details. Thus, for the case in treatment, this gives rise to the so called affine geometry of surfaces. For a given surface in the three-dimensional space we use, within this context, concepts such as "affine normal" and "affine distance", corresponding to the above mentioned ones in Euclidean geometry. See the description of the corresponding theory as exposed, in a much wider sense, in Gigena (1993, 1996a, b) and, with somehow different kind of notation, also in Nomizu and Sasaki (1994).

We introduce, in Section 2 of the present article, an abbreviated version of the concept of Affine Shell, already developed in the previously cited articles. The treatment of Compatibility Conditions occupies Section 3, while the Basic Inequalities of the Theory are considered in Section 4. The further development of the Theory consists in the presentation of the Strain-Stress Relations in Affine Shells which is taken care of in Section 5. Section 6 is devoted to the treatment of the Estimates for the $L_{2}$-Norms of Second Order and, finally, in section 7 we come to conclude this article by exposing the estimates of Higher Order Derivatives, both partial and covariant.

## 2 AN ABBREVIATED VERSION OF AFFINE SHELLS

We consider the middle surface of a (solid) shell in its original (undeformed) state, denoted by $M_{0}$, parametrized locally by a vector function $X_{0}: U \rightarrow \mathbb{R}^{3}$, where $U \subset \mathbb{R}^{2}$, which is assumed to be enough smooth. Coordinates in the domain are denoted by $\left(u^{1}, u^{2}\right)$. Thus, we can write locally $M_{0}=X_{0}\left(u^{1}, u^{2}\right)$ and assume besides, as it is usually understood, that $X_{0}$ is a topological immersion (embedding). Particles in the original state have curvilinear Lagrange coordinates $\left(U^{1}, U^{2}, U^{3}\right)$ that for our present purposes shall be chosen in a special way, by writing: $\left(U^{1}, U^{2}, U^{3}\right)=\left(u^{1}, u^{2}, u\right), X\left(u^{1}, u^{2}, u\right)=X_{u}\left(u^{1}, u^{2}\right)=X_{0}\left(u^{1}, u^{2}\right)+u \overrightarrow{\mathbf{n}}$, where we have obviously extended the previous function to $X: U \times(-h, h) \rightarrow \mathbb{R}^{3}$, and $\overrightarrow{\mathbf{n}}$ is the vector field normal to the middle surface. This normal can be the Euclidean normal, $N_{e u}$, of the classical, Euclidean Theory of Surfaces, or the Unimodular Affine normal, $N_{u a}$, of our own, current development. In each case, we shall clarify the situation when we deal with one
or the other.
In the Euclidean case we shall use the following notations regarding the main geometrical objects, defined on the middle surface prior to deformation, that take part in the formulation of the theory, as treated mainly in Gigena et al. (2002, 2003, 2004, 2005); John (1965 and 1971); Millman and Parker (1977).

$$
\begin{equation*}
I_{e u}=\sum_{\alpha, \beta} a_{\alpha \beta} d u^{\alpha} d u^{\beta} \quad \text { with } \quad a_{\alpha \beta}=\frac{\partial X_{0}}{\partial u^{\alpha}} \cdot \frac{\partial X_{0}}{\partial u^{\beta}} \tag{1}
\end{equation*}
$$

denotes the Euclidean first fundamental form, while with the expression

$$
\begin{equation*}
I I_{e u}=\sum_{\alpha, \beta} L_{\alpha \beta} d u^{\alpha} d u^{\beta} \quad \text { where } \quad L_{\alpha \beta}=N_{e u} \cdot \frac{\partial^{2} X_{0}}{\partial u^{\beta} \partial u^{\alpha}}, \tag{2}
\end{equation*}
$$

we represent the second fundamental form, and with

$$
\begin{equation*}
I I I_{e u}=\sum_{\alpha, \beta} M_{\alpha \beta} d u^{\alpha} d u^{\beta}, \text { where } \quad M_{\alpha \beta}=\sum_{\lambda} L_{\alpha \lambda} L_{\beta}^{\lambda}=\sum_{\eta \lambda} a^{\nu \lambda} L_{\alpha \lambda} L_{\beta \gamma}, \tag{3}
\end{equation*}
$$

the Euclidean third fundamental form.
In the state previous to deformation the border of the shell is made up of two "faces", which are surfaces parallel to the middle surface $M_{0}$ at respective distance $h$, measured along the Euclidean normal $N_{e u}$, and of the "border" constituted by segments normal to the faces. Therefore, along the normal to $M_{0}$ coordinates $U^{1}, U^{2}$ remain constant while $U^{3}:=u$ measures the signed distance from $M_{0}$. Faces can be represented, then, by equations $U^{3}=u= \pm h$ while the middle surface is given by $U^{3}=u=0$.

Now if $a_{\alpha \beta}, L_{\alpha \beta}, M_{\alpha \beta}$, are respectively the coefficients of the first, second and third Euclidean fundamental forms of the middle surface $M_{0}$, the Euclidean structure of the ambient space induces a Riemannian structure on the shell and we can obtain, by means of a straightforward computation, the following expressions in normal coordinates $\left(U^{1}, U^{2}, U^{3}\right)=\left(u^{1}, u^{2}, u\right):$

$$
\begin{align*}
& A_{\alpha \beta}=\frac{\partial X}{\partial u^{\alpha}} \cdot \frac{\partial X}{\partial u^{\beta}}=a_{\alpha \beta}-2 u L_{\alpha \beta}+u^{2} M_{\alpha \beta}  \tag{4}\\
& A_{\alpha 3}=A_{3 \alpha}=\frac{\partial X}{\partial u^{\alpha}} \cdot \frac{\partial X}{\partial t}=\frac{\partial X}{\partial u^{\alpha}} \cdot N_{e u}=0,  \tag{5}\\
& A_{33}=\frac{\partial X}{\partial t} \cdot \frac{\partial X}{\partial t}=N_{e u} \cdot N_{e u}=1 . \tag{6}
\end{align*}
$$

Corresponding to the shell, and its middle surface, in the state previous to deformation, we can consider the geometrical objects belonging to the shell in the deformed state that we shall denote with an upper right asterisk. Thus, for example, $X_{0}^{*}: U \rightarrow \mathbb{R}^{3}$, where $U \subset \mathbb{R}^{2}$, represents the parametrization of the deformed middle surface $M_{0}^{*}=X_{0}^{*}\left(u^{1}, u^{2}\right)$, and we remark that the domain of definition of this immersion, $U \subset \mathbb{R}^{2}$, and the parameters ( $u^{1}, u^{2}$ ) used in it, are the same as those belonging to the middle surface of the shell in the original state, previous to deformation.

Consequently, the rest of geometrical objects change from one state to the other and the
problem is to determine the nature and extension of such changes for every one of them reducing, under appropriate hypotheses, the obtainable information to both middle surfaces. One such hypothesis is the one concerning the comparison of the thickness parameter $h$, which it is usually assumed to be small with respect to the other dimensions of the shell. This introduces in the theory the concept of "thin" shell which has important uses and applications.

Considering now the Unimodular Affine Geometry of Surfaces, we need to assume defined, in the ambient space $\mathbb{R}^{3}$ an exterior 3-form, or non-trivial determinant function, denoted by the symbol $[,]=,\operatorname{det}$. Then, given the same previous mean surface, we represent the objects of that geometry by the following expressions:

In order to construct the Unimodular first fundamental form we define, firstly

$$
\begin{equation*}
h_{\alpha \beta}=\left[\frac{\partial X_{0}}{\partial u^{1}}, \frac{\partial X_{0}}{\partial u^{2}}, \frac{\partial^{2} X_{0}}{\partial u^{\alpha} \partial u^{\beta}}\right], \tag{7}
\end{equation*}
$$

then, if we assume that the surface is non-degenerate, i.e., $H=\operatorname{det}\left(h_{\alpha \beta}\right) \neq 0$, we can write $g_{\alpha \beta}=|H|^{-1 / 4} h_{\alpha \beta}$, obtaining the Unimodular Affine First Fundamental Form expressed by equation

$$
\begin{equation*}
I_{u a}=\sum_{\alpha, \beta} g_{\alpha \beta} d u^{\alpha} d u^{\beta}, \tag{8}
\end{equation*}
$$

that turns out to be a semi-Riemannian structure, Gigena (1993, 1996a, b); Nomizu and Sasaki (1994). The Unimodular Affine Normal is defined now by the expression

$$
\begin{equation*}
N_{u a}=\frac{1}{2} \Delta\left(X_{0}\right), \tag{9}
\end{equation*}
$$

where $\Delta$ is the Laplacian operator with respect to the pseudometric $I_{u a}$, i.e.:

$$
\begin{equation*}
\Delta X_{0}=\frac{1}{\sqrt{|g|}} \sum_{\alpha=1}^{2} \frac{\partial}{\partial u^{\alpha}}\left(\sqrt{|g|} \sum_{\beta=1}^{2} g^{\alpha \beta} \frac{\partial X_{0}}{\partial u^{\beta}}\right) \text { with } g=\operatorname{det}\left(g_{\alpha \beta}\right) \tag{10}
\end{equation*}
$$

From the above we obtain three connections:

1) The Levi-Civita connection with respect to the Euclidean metric $I_{e u}$, that we shall label here as $\nabla_{e u}$ and which coincides with the projection over $M_{0}$ of the usual, flat connection $D$ of $\mathbb{R}^{3}$ in the direction of the classical Euclidean normal $N_{e u}$.
2) The Levi-Civita connection with respect to pseudometric $I_{u a}: \tilde{\nabla}$.
3) The affine normal induced connection: $\nabla$, i.e., the projection of $D$ in the direction of $N_{\text {ua }}$ :

$$
\begin{equation*}
\nabla_{X_{p}} Y=\operatorname{proy}_{N_{u a}}\left(D_{X_{p}} Y\right) . \tag{11}
\end{equation*}
$$

We define next the Unimodular Affine Second Fundamental Form, as previously introduced in Gigena (1993, 1996a, b):

$$
\begin{equation*}
\nabla\left(I_{u a}\right)=I I_{u a}, \tag{12}
\end{equation*}
$$

that we also represent in local coordinates by:

$$
\begin{equation*}
I I_{u a}=\sum_{\alpha \beta r} g_{\alpha \beta \gamma} d u^{\alpha} d u^{\beta} d u^{\gamma}, \tag{13}
\end{equation*}
$$

with the coefficients $g_{\alpha \beta \gamma}$ totally symmetric in their indices. Some authors prefer to refer to the latter as the Cubic Form, see Nomizu and Sasaki (1994).

Finally, we consider the Affine Third Fundamental Form that we can describe in the following way: similar to the Euclidean case regarding the Weingarten equation, it turns out too in affine geometry of surfaces that the local derivatives of the affine normal belong to the tangent plane of the surface at each point, i.e., we can write

$$
\begin{equation*}
\frac{\partial N_{u a}}{\partial u^{\alpha}}=-\sum_{\beta} B_{\alpha}^{\beta} \frac{\partial X_{0}}{\partial u^{\beta}}=-B_{\alpha}^{1} \frac{\partial X_{0}}{\partial u^{1}}-B_{\alpha}^{2} \frac{\partial X_{0}}{\partial u^{2}}, \tag{14}
\end{equation*}
$$

and define the Affine Third Fundamental Form by the expression:

$$
\begin{equation*}
I I I_{u a}=B_{\alpha \beta} d u^{\alpha} d u^{\beta} \quad \text { with } \quad B_{\alpha \beta}=\sum_{\gamma} g_{\alpha \gamma} B_{\beta}^{\gamma} . \tag{15}
\end{equation*}
$$

As we have previously seen, the definition of shell as a three-dimensional body and, in particular, the Riemannian structure induced on that object by the ambient space metric is generated in a natural fashion. In the present case of Unimodular Affine Geometry that extension is not at all that immediate. However, as we shall see, it can also be realized in a canonical way. We start from the affine invariant pseudometric $I_{u a}$, defined on the middle surface $M_{0}$ :

$$
\begin{equation*}
g_{\alpha \beta}=I_{\text {иа }}\left(\frac{\partial X_{0}}{\partial u^{\alpha}}, \frac{\partial X_{0}}{\partial u^{\beta}}\right) . \tag{16}
\end{equation*}
$$

In the present context we define on the shell a pseudo-metric, which is a Unimodular Affine invariant, to be denoted by

$$
\begin{equation*}
G=\sum G_{i j} d u^{i} d u^{j}, \tag{17}
\end{equation*}
$$

i.e., with $G_{i j}:=G\left(\frac{\partial X}{\partial u^{i}}, \frac{\partial X}{\partial u^{j}}\right)$.

Let us observe first of all that, since bilinearity must be preserved, we have to write in affine normal coordinates of the shell

$$
\begin{align*}
G_{\alpha \beta} & =G\left(\frac{\partial X}{\partial u^{\alpha}}, \frac{\partial X}{\partial u^{\beta}}\right) \\
& =G\left(\frac{\partial X_{0}}{\partial u^{\alpha}}+u \frac{\partial N_{\mathrm{ua}}}{\partial u^{\alpha}}, \frac{\partial X_{0}}{\partial u^{\beta}}+u \frac{\partial N_{\mathrm{ua}}}{\partial u^{\beta}}\right) \tag{18}
\end{align*}
$$

where, by definition

$$
\begin{equation*}
G_{\alpha \beta}:=g_{\alpha \beta}-2 u B_{\alpha \beta}+u^{2} \sum_{\lambda} B_{\alpha}^{\lambda} B_{\beta \lambda} \tag{19}
\end{equation*}
$$

and where, as stated previously, Greek indices run from 1 to 2 . Thus, in order to extend that definition to the third index, we also write:

$$
\begin{equation*}
G_{3 \alpha}=G_{\alpha 3}=G\left(\frac{\partial X}{\partial u^{\alpha}}, \frac{\partial X}{\partial u^{3}}\right)=G\left(X_{\alpha}, N_{u a}\right):=0 \tag{20}
\end{equation*}
$$

and, finally,

$$
\begin{equation*}
G_{33}=G\left(\frac{\partial X}{\partial u^{3}}, \frac{\partial X}{\partial u^{3}}\right)=G\left(N_{u a}, N_{u a}\right):=1 . \tag{21}
\end{equation*}
$$

It is easy to see that, for $u=u^{3}$ enough small, it holds:

$$
\begin{equation*}
\operatorname{det}\left(G_{i j}\right) \neq 0 \tag{22}
\end{equation*}
$$

and, consequently, the latter is a pseudo-Riemannian, Unimodular affine invariant metric defined on the shell, as it was our purpose to construct.

## 3 COMPATIBILITY CONDITIONS

One of the main aspects in the theory of shells is the determination of compatibility conditions. These are conditions obtained on the behavior of the various difference tensors that can be defined by comparing the two states of the shell. The natural tool here is represented by the integrability conditions that must be satisfied, in all cases, by both middle surfaces. These conditions are very well known in the case of Euclidean shells, see, for example John (1965 and 1971), Koiter (1970), Möllmann (1981), and can be described, in our present own notation, as follows:

For the tensor with components defined by $\varepsilon_{\alpha \beta}=\frac{1}{2}\left(a_{\alpha \beta}^{*}-a_{\alpha \beta}\right)$, it is proven that

$$
\begin{equation*}
\varepsilon_{\beta, \delta}^{\beta, \delta}-\varepsilon_{\beta, \delta}^{\delta, \beta}=L_{\beta}^{* \delta} L_{\delta}^{* \beta}-L_{\delta}^{* \delta} L_{\beta}^{* \beta}+g_{\mu}^{* \beta}\left(L_{\delta}^{\delta} L_{\beta}^{\mu}-L_{\beta}^{\delta} L_{\delta}^{\mu}\right)-g_{\mu \nu}^{*} g^{\beta \alpha} g^{\delta \gamma}\left(C_{\alpha \beta}^{\mu} C_{\gamma \delta}^{v}-C_{\alpha \delta}^{\mu} C_{\beta \gamma}^{v}\right) \tag{23}
\end{equation*}
$$

while for the difference tensor $w_{\alpha \beta}:=L_{\alpha \beta}^{*}-L_{\alpha \beta}$ it holds

$$
\begin{equation*}
w_{\beta, \gamma}^{\alpha}-w_{\gamma, \beta}^{\alpha}=g^{\alpha \rho}\left(w_{\rho \beta, \gamma}-w_{\rho \gamma, \beta}\right)=g^{\alpha \rho}\left(C_{\rho \gamma}^{\mu} L_{\mu \beta}^{*}-C_{\rho \beta}^{\mu} L_{\mu \gamma}^{*}\right) . \tag{24}
\end{equation*}
$$

In both equations the symbol $C_{\rho \beta}^{\mu}$ represent the components of the difference tensor between the Levi-Civita connection of $M_{0}^{*}$ and that of $M_{0}$.

Now, for the case of affine shells the corresponding compatibility conditions were obtained in our previous article Gigena et al. (2002), and can be summarized as follows.

For the difference tensors defined by the various expressions that establish comparisons between the first, second and third fundamental forms, i.e.,

$$
\begin{equation*}
\varepsilon_{\alpha \beta}=\frac{1}{2}\left(g_{\alpha \beta}^{*}-g_{\alpha \beta}\right), \quad \sigma_{\alpha \beta \gamma}:=g_{\alpha \beta \gamma}^{*}-g_{\alpha \beta \gamma}, \quad w_{\alpha \beta}:=B_{\alpha \beta}^{*}-B_{\alpha \beta}, \tag{25}
\end{equation*}
$$

and the tensor defined by comparison between the corresponding Levi-Civita connections, represented by equation: $\tilde{\Gamma}_{\alpha \beta}^{*} \mu=C_{\alpha \beta}^{\mu}+\tilde{\Gamma}_{\alpha \beta}^{\mu}$ there hold the following conditions:

## 1) Affine Gauss condition

$$
\begin{align*}
\varepsilon_{\beta, \delta}^{\beta, \delta}-\varepsilon_{\beta, \delta}^{\delta, \beta} & =\frac{1}{2}\left(B_{\beta}^{* \beta} g_{\delta}^{* \delta}-B_{\delta}^{* \beta} g_{\beta}^{* \delta}+B_{\delta}^{* \delta} g_{\beta}^{* \beta}-B_{\beta}^{* \delta} g_{\delta}^{* \beta}\right)-\frac{1}{2} g^{\beta \alpha} g^{\delta \gamma}\left(A_{\gamma \beta}^{* \eta} \cdot g_{\alpha \eta \delta}^{*}-A_{\gamma \delta}^{* \eta} \cdot g_{\alpha \eta \beta}^{*}\right)-  \tag{26}\\
& +\frac{1}{2} g_{\mu}^{* \beta}\left(-B_{\delta}^{\delta} \delta_{\beta}^{\mu}+A_{\beta}^{\delta \eta} \cdot A_{\eta \delta}^{\mu}\right)-g_{\lambda \mu}^{*} g^{\beta \alpha} g^{\delta \gamma}\left(C_{\alpha \beta}^{\lambda} C_{\gamma \delta}^{\mu}-C_{\alpha \delta}^{\lambda} C_{\beta \gamma}^{\mu}\right)
\end{align*}
$$

2) Affine Mainardi-Codazzi condition

$$
\begin{align*}
\sigma_{\alpha \beta \gamma, \delta}-\sigma_{\alpha \beta \delta, \gamma} & =g_{\mu \beta \gamma}^{*} C_{\alpha \delta}^{\mu}+g_{\mu \alpha \gamma}^{*} C_{\beta \delta}^{\mu}-g_{\mu \beta \delta}^{*} C_{\alpha \gamma}^{\mu}-g_{\mu \alpha \delta}^{*} C_{\beta \gamma}^{\mu}+ \\
& +B_{\alpha \delta}^{*} g_{\beta \gamma}^{*}+B_{\beta \delta}^{*} g_{\alpha \gamma}^{*}-B_{\alpha \gamma}^{*} g_{\beta \delta}^{*}-B_{\beta \gamma}^{*} g_{\alpha \delta}^{*}-  \tag{27}\\
& -B_{\alpha \delta} g_{\beta \gamma}-B_{\beta \delta} g_{\alpha \gamma}+B_{\alpha \gamma} g_{\beta \delta}+B_{\beta \gamma} g_{\alpha \delta}
\end{align*}
$$

3) Codazzi condition for the affine shape operators

$$
\begin{equation*}
w_{\beta, \alpha}^{\alpha}-w_{\alpha, \beta}^{\alpha}=g^{\alpha \rho}\left[B_{\beta \mu}^{*}\left(C_{\rho \alpha}^{\mu}+A_{\rho \alpha}^{* \mu}\right)-B_{\alpha \mu}^{*}\left(C_{\rho \lambda}^{\mu}+A_{\rho \beta}^{* \mu}\right)+B_{\alpha \mu} A_{\rho \beta}^{\mu}\right] . \tag{28}
\end{equation*}
$$

## 4 BASIC INEQUALITIES FOR AFFINE SHELLS

The following basic inequalities, involving the geometrical objects treated before, were previously obtained in Gigena et al. (2004). When represented in the form of Monge's, i.e., as a graph, the middle surface of the shell $M_{0}$ has all of its geometrical properties related to a given function $f$ assumed to be enough differentiable and, in the present context of affine geometry, satisfying a partial differential equation of Monge-Ampère type:

$$
\begin{equation*}
\operatorname{det}\left(\partial_{\alpha \beta} f\right)= \pm F \tag{29}
\end{equation*}
$$

and for such a kind of equations, with boundary conditions as in the present case, there hold bounds for the function $f$ and its derivatives. Also, since the function $F$ is strictly positive in the (compact) domain where $f$ is defined, there exist lower an upper bounds for $F$ as well. It is very convenient, in order to avoid unwanted complications in notation, that the lower bound, which is otherwise positive, be taken to be exactly equal to +1 , by means of a suitable rescaling if it were necessary.

As a consequence, we can also assume that the second derivatives of $f$ are bounded, i.e., the components of the Hessian matrix $\left(\partial_{\alpha \beta} f\right)$, the components of the inverse matrix of the latter, denoted by $\left(f^{\alpha \beta}\right)$, and the components of the pseudometric tensor, covariant as well as contravariant, i.e., $g_{\alpha \beta}$ and $g^{\alpha \beta}$. These facts being expressed in the following inequalities:

$$
\begin{equation*}
\left|\partial_{\alpha \beta} f\right|<K,\left|f^{\alpha \beta}\right|<K, \quad\left|g_{\alpha \beta}\right|<K, \quad\left|g^{\alpha \beta}\right|<K . \tag{30}
\end{equation*}
$$

Besides, since the higher order derivatives are also bounded, and in order to unify notation, we shall assume that there exists a generalized affine upper bound of curvature, intimately related to the upper bound for the affine principal curvatures of the middle surface $M_{0}$, that we shall also denote by $R$, and that for the present, affine case, remains specified by the conditions that:

$$
\begin{equation*}
\left|\partial_{\alpha \beta \gamma} f\right|<\frac{1}{R^{1 / 2}}, \tag{31}
\end{equation*}
$$

and for the successive derivatives,

$$
\begin{gather*}
\left|\partial_{\alpha \beta \eta} f\right| \leq R^{-1},  \tag{32}\\
\left|\partial_{\alpha \beta \eta \eta} f\right| \leq\left(R^{-1 / 2}\right)^{3} ; \ldots \tag{3}
\end{gather*}
$$

be satisfied for as high order of derivatives as needed in the development of the theory.
By using these hypotheses one obtains the corresponding bounds for the components of the tensor representing the third fundamental form:

$$
\begin{equation*}
B_{\alpha \beta}=-\frac{1}{4}\left(\partial_{\alpha \beta}(\log F)+\frac{1}{4} \partial_{\alpha}(\log F) \partial_{\beta}(\log F)-\sum_{\sigma, \lambda} f^{\sigma \lambda} \partial_{\alpha \beta \sigma} f \partial_{\lambda}(\log F)\right), \tag{34}
\end{equation*}
$$

if we have in mind, besides, the two following, well-known identities:

$$
\begin{gather*}
\partial_{\alpha} \log F=\sum_{\rho, \sigma} f^{\rho \sigma} \partial_{\alpha \rho \sigma} f,  \tag{35}\\
\partial_{\alpha \beta} \log F=\sum_{\rho, \sigma} f^{\rho \sigma} \partial_{\alpha \beta \rho \sigma} f-\sum_{\rho, \sigma} f^{\rho \theta} f^{\tau \sigma} \partial_{\theta \tau \alpha} f \partial_{\rho \sigma \beta} f, \tag{36}
\end{gather*}
$$

with which it turns out that:

$$
\begin{equation*}
\left|B_{\alpha \beta}\right| \leq \frac{1}{4} \frac{1}{R}\left(4 K+32 K^{2}\right)=\left(K+8 K^{2}\right) \frac{1}{R} . \tag{37}
\end{equation*}
$$

We compute next the partial derivatives of these components

$$
\begin{align*}
\partial_{\gamma} B_{\alpha \beta}= & -\frac{1}{4}\left(\begin{array}{l}
\partial_{\gamma} f^{\rho \tau} \partial_{\alpha \beta \rho \tau} f+f^{\rho \tau} \partial_{\alpha \beta \rho \tau \gamma} f-\left(\partial_{\gamma} f^{\rho \eta} \partial_{\eta \mu \alpha} f\right)\left(f^{\mu \tau} \partial_{\beta \rho \tau} f\right)- \\
-\left(f^{\rho \eta} \partial_{\eta \mu \alpha \gamma} f\right)\left(f^{\mu \tau} \partial_{\beta \rho \tau} f\right)-\left(f^{\rho \eta} \partial_{\eta \mu \alpha} f\right)\left(\partial_{\gamma} f^{\mu \tau} \partial_{\beta \rho \tau} f\right)- \\
-\left(f^{\rho \eta} \partial_{\eta \mu \alpha} f\right)\left(f^{\mu \tau} \partial_{\beta \rho \tau \gamma} f\right)+
\end{array}\right)+ \\
& +\frac{1}{16}\binom{\left(\partial_{\gamma} f^{\rho \tau} \partial_{\tau \rho \alpha} f\right)\left(f^{\sigma \tau} \partial_{\beta \sigma \lambda} f\right)+\left(f^{\rho \tau} \partial_{\tau \rho \alpha \gamma} f\right)\left(f^{\sigma \lambda} \partial_{\beta \sigma \lambda} f\right)+}{\left(f^{\rho \tau} \partial_{\tau \rho \alpha} f\right)\left(\partial_{\gamma} \sigma^{\sigma \lambda} \partial_{\beta \sigma \lambda} f\right)+\left(f^{\rho \tau} \partial_{\tau \rho \alpha} f\right)\left(f^{\sigma \lambda} \partial_{\beta \sigma \lambda \lambda} f\right)}-  \tag{38}\\
& -\frac{1}{4}\binom{\left(\partial_{\gamma} f^{\sigma \lambda} \partial_{\alpha \beta \sigma} f\right)\left(f^{\rho \tau} \partial_{\lambda \rho \tau} f\right)+\left(f^{\sigma \lambda} \partial_{\alpha \beta \sigma \gamma} f\right)\left(f^{\rho \tau} \partial_{\lambda \rho \tau} f\right)+}{+\left(f^{\sigma \lambda} \partial_{\alpha \beta \sigma} f\right)\left(\partial_{\gamma} f^{\rho \tau} \partial_{\lambda \rho \tau} f\right)+\left(f^{\sigma \lambda} \partial_{\alpha \beta \sigma} f\right)\left(f^{\rho \tau} \partial_{\lambda \rho \tau \gamma} f\right)} .
\end{align*}
$$

Then, by using the identity $\sum_{\lambda} f^{\lambda \sigma} \partial_{\lambda \mu} f=\delta_{\mu}^{\sigma}$, from which it follows that

$$
\begin{equation*}
\partial_{\alpha} f^{\sigma \gamma}=-\sum_{\lambda} f^{\lambda \gamma} f^{\mu \sigma} \partial_{\alpha \chi \mu} f, \tag{39}
\end{equation*}
$$

we find by direct computation the following estimate

$$
\begin{equation*}
\left|\partial_{\gamma} B_{\alpha \beta}\right| \leq \frac{1}{R^{3 / 2}}\left(K+19 K^{2}+24 K^{3}\right) . \tag{40}
\end{equation*}
$$

With the development done so far, we can also obtain estimates for the components of the pseudometric, i.e., the components of the pseudo-Riemannian tensor $G=\sum_{i, j} G_{i j} d u^{i} d u^{j}$ of the shell in the undeformed state, and its successive derivatives, partial as well as covariant. For example, from

$$
\begin{equation*}
G_{\alpha \beta}:=g_{\alpha \beta}-2 u B_{\alpha \beta}+u^{2} \sum_{\lambda} B_{\alpha}^{\lambda} B_{\beta \lambda}, \tag{41}
\end{equation*}
$$

we obtain, firstly, that

$$
\begin{equation*}
G_{\alpha \beta}=g_{\alpha \beta}-2 u B_{\alpha \beta}+u^{2} \sum_{\lambda, \mu} g^{\lambda \mu} B_{\alpha \mu} B_{\beta \lambda}, \tag{42}
\end{equation*}
$$

and, consequently

$$
\begin{equation*}
\left|G_{\alpha \beta}\right| \leq K+2 K(1+8 K) \frac{h}{R}+4 K^{2}(1+8 K)\left(\frac{h}{R}\right)^{2} . \tag{43}
\end{equation*}
$$

## 5 STRAIN-STRESS RELATIONS IN AFFINE SHELLS

We recall that, for the present case of affine shells, the contravariant components of the stress tensor, $t^{m k}$, are connected with the components of the strain tensor, $\varepsilon_{m k}$, by means of the stress-strain relations

$$
\begin{equation*}
t^{m k}:=\sqrt{\frac{G}{G^{*}}} \frac{\partial W}{\partial \varepsilon_{m k}}, \tag{44}
\end{equation*}
$$

defined in a similar fashion as to the Euclidean case, introduced by John (1965), where $W$ is the strain energy density of the given material.

The same expression, in terms of the corresponding (1,1)-tensors is

$$
\begin{equation*}
t_{i}^{m}=\sum_{k} G_{i k} t^{m k}=\sqrt{\frac{G}{G^{*}}} \frac{\partial W}{\partial \varepsilon_{m}^{i}} . \tag{45}
\end{equation*}
$$

Next, we introduce the components of the "pseudo-stress tensor" defined by

$$
\begin{equation*}
T_{j}^{m}:=\sqrt{\frac{G^{*}}{G}} t_{j}^{m}-\delta_{j}^{m} W, \tag{46}
\end{equation*}
$$

and we may also write

$$
\begin{equation*}
T_{i}^{m}=\left(W_{s_{1}}-W\right) \delta_{i}^{m}+\left(2 W_{s_{1}}+2 W_{s_{2}}\right) \varepsilon_{i}^{m}+\left(4 W_{s_{2}}+3 W_{s_{3}}\right) \sum_{k} \varepsilon_{i}^{k} \varepsilon_{k}^{m}+6 W_{s_{3}} \sum_{s, k} \varepsilon_{i}^{s} \varepsilon_{s}^{k} \varepsilon_{k}^{m} \tag{47}
\end{equation*}
$$

where

$$
\begin{equation*}
s_{1}=\sum_{i} \varepsilon_{i}^{i}, \quad s_{2}=\sum_{i, j} \varepsilon_{j}^{i} \varepsilon_{i}^{j}, \quad s_{3}=\sum_{i, j, k} \varepsilon_{j}^{i} \varepsilon_{k}^{j} \varepsilon_{i}^{k} \tag{48}
\end{equation*}
$$

Then, the equations of equilibrium can be written

$$
\begin{equation*}
t^{i j}{ }_{, j}+c_{h j}^{i} t^{h j}+c_{h j}^{h} t^{i j}=0, \tag{49}
\end{equation*}
$$

where

$$
\begin{equation*}
c_{j k}^{i}=\frac{1}{2} \bar{G}^{* i r}\left(G_{r j, k}^{*}+G_{r k,, j}^{*}-G_{j k, r}^{*}\right) \tag{50}
\end{equation*}
$$

and where we also have, as a consequence, that

$$
\begin{equation*}
\sum_{m} T_{i ; m}^{m}=\left(\sqrt{\frac{G^{*}}{G}}\right) \sum_{m, r, s}\left(c_{m r}^{r} t^{m s} G_{s i}^{*}-t^{r s} c_{r m}^{m} G_{s i}^{*}-t^{m r} c_{r m}^{s} G_{s i}^{*}+t^{m s} G_{s i ; m}^{*}-\frac{1}{2} t^{m s} G_{s m ; i}^{*}\right)=0 . \tag{51}
\end{equation*}
$$

Additional notations are needed in order to compare components of stress and strain tensors, even those belonging to different spaces of definition. For example, and very particularly, in order to compare components of type $(0,2)$ tensors $t^{i j}$, with those with components of type $(1,1) t_{i}^{j}$. Thus, we follow in this respect the kind of notation previously introduced by Fritz John, see John (1965). In particular the so-called "general form of an expression" like

$$
\begin{equation*}
F(p, q)(u+v+w), \tag{52}
\end{equation*}
$$

representing a vector, in a suitable space, where $u, v, w, p, q$ are vectors themselves. The notation indicates that each of the components of $F(p, q)(u+v+\omega)$ is a sum of a linear form in the components of $u$, a linear form in the components of $v$, and a linear form in the components of $w$. The coefficients of these linear forms are functions of the components of the vectors $p$ and $q$ defined and differentiable as often as needed for all sufficiently small "lengths" $|p|$ and $|q|$. The letter $F$ stands for a different expression in every equation to be considered. Thus, for example, we can write, for the components of the stress and strain tensors, of type $(1,1)$

$$
\begin{equation*}
t=\left(t_{k}^{i}\right) \text { and } \varepsilon=\left(\varepsilon_{k}^{i}\right), \tag{53}
\end{equation*}
$$

and in terms of the Lamé coefficients $\lambda, \mu$, the following equation

$$
\begin{equation*}
t_{i}^{m}=\lambda \sum_{j} \varepsilon_{j}^{j} \delta_{i}^{m}+2 \mu \varepsilon_{i}^{m}+F(\varepsilon) \varepsilon^{2} \tag{54}
\end{equation*}
$$

since such coefficients are defined by the relation

$$
\begin{equation*}
W:=\frac{\lambda}{2}\left(s_{1}\right)^{2}+\mu s_{2}+F(\varepsilon) \varepsilon^{3} . \tag{55}
\end{equation*}
$$

where, $s_{1}=\sum_{i} \varepsilon_{i}^{i}, s_{2}=\sum_{i, j} \varepsilon_{j}^{i} \varepsilon_{i}^{j}, s_{3}=\sum_{i, j, k} \varepsilon_{j}^{i} \varepsilon_{k}^{j} \varepsilon_{i}^{k}$, as in (48), and where we observe that the first two terms, on the right-hand side, are quadratic in terms of the strain tensor (operator) $\varepsilon=\left(\varepsilon_{k}^{i}\right)$, while the third term involves all of those components of order higher than two, representing otherwise the "remainder", of paramount importance when coming to the
corresponding numerical estimates.
From now on we establish that in the same sense have to be interpreted all of the expressions to follow. Hence, by taking partial derivatives, we can write

$$
\begin{equation*}
W_{s_{1}}=\frac{\partial W}{\partial s_{1}}=\partial_{s_{1}} W=\lambda s_{1}, W_{s_{2}}=\frac{\partial W}{\partial s_{2}}=\partial_{s_{2}} W=\mu . \tag{56}
\end{equation*}
$$

From the latter we obtain, successively:

$$
\begin{gather*}
2 W_{s_{1}}+2 W_{s_{2}}=2 \mu+2 \lambda s_{1},  \tag{57}\\
4 W_{s_{2}}+3 W_{s_{3}}=2 \mu+F(t) t^{3},  \tag{58}\\
W_{s_{1}}-W=\lambda s_{1}-\frac{1}{2} \lambda\left(s_{1}\right)^{2}-\frac{1}{2} \mu s_{2}+F(\varepsilon) \varepsilon^{3} . \tag{59}
\end{gather*}
$$

Then, by using the Taylor's series development

$$
\begin{equation*}
(1+x)^{-1 / 2}=1+\left(-\frac{1}{2}\right) x+\frac{3 / 4}{2} x^{2}+\ldots \tag{60}
\end{equation*}
$$

we can express

$$
\begin{equation*}
\sqrt{\frac{G}{G^{*}}}=\left(\frac{G}{G^{*}}\right)^{1 / 2}=\left(\frac{G^{*}}{G}\right)^{-1 / 2}=1-\frac{1}{2} s_{1}+\frac{3}{8}\left(s_{1}\right)^{2}+\ldots \tag{61}
\end{equation*}
$$

and

$$
\begin{equation*}
t_{i}^{m}=\sqrt{\frac{G}{G^{*}}}\left(W_{s_{1}} \delta_{i}^{m}+2 W_{s_{2}} \varepsilon_{i}^{m}+3 W_{s_{3}} \sum_{j} \varepsilon_{j}^{m} \varepsilon_{i}^{j}\right) \tag{62}
\end{equation*}
$$

becomes, first

$$
\begin{equation*}
t_{i}^{m}=\left(1-\frac{1}{2} s_{1}+\frac{3}{8}\left(s_{1}\right)^{2}+\ldots\right)\left(W_{s_{1}} \delta_{i}^{m}+2 W_{s_{2}} \varepsilon_{i}^{m}+3 W_{s_{3}} \sum_{j} \varepsilon_{j}^{m} \varepsilon_{i}^{j}\right) \tag{63}
\end{equation*}
$$

and, afterwards

$$
\begin{align*}
t_{i}^{m} & =W_{s_{1}} \delta_{i}^{m}+2 W_{s_{2}} \varepsilon_{i}^{m}+3 W_{s_{3}} \sum_{j} \varepsilon_{j}^{m} \varepsilon_{i}^{j} \\
& =\frac{\lambda}{2} s_{1} \delta_{i}^{m}+2 \mu \varepsilon_{i}^{m}+3 W_{s_{3}} \sum_{j} \varepsilon_{j}^{m} \varepsilon_{i}^{j}  \tag{64}\\
& =\frac{\lambda}{2} \sum_{j} \varepsilon_{j}^{j} \delta_{i}^{m}+2 \mu \varepsilon_{i}^{m}+F(\varepsilon) \varepsilon^{2} .
\end{align*}
$$

From the latter, the trace of the stress tensor (operator) can be written

$$
\begin{equation*}
\sum_{j} t_{j}^{j}=\left(\frac{3}{2} \lambda+2 \mu\right) \sum_{j} \varepsilon_{j}^{j}+F(\varepsilon) \varepsilon^{2} \tag{65}
\end{equation*}
$$

where $\sum_{j} \varepsilon_{j}^{j}$ itself represents the trace of the strain tensor (operator).
Hence, from the above we can also write

$$
\begin{equation*}
\varepsilon_{i}^{m}=\frac{1}{2 \mu} t_{i}^{m}-\frac{\lambda}{2 \mu} s_{1} \delta_{i}^{m}+F(t) t^{2} \tag{66}
\end{equation*}
$$

or, also,

$$
\begin{equation*}
\varepsilon_{i}^{m}=\frac{1}{2 \mu} t_{i}^{m}-\frac{1-2 \mu}{2 \mu} \sum_{j} t_{j}^{j} \delta_{i}^{m}+F(t) t^{2} . \tag{67}
\end{equation*}
$$

Then, the expression for the components of the pseudo-stress tensor is

$$
\begin{align*}
T_{i}^{m}=t_{i}^{m} & +\frac{2}{\mu} \sum_{i} t_{i}^{m} t_{i}^{s}+\left(\frac{5 \mu-2}{\mu}\right) \sum_{j} t_{j}^{j} t_{i}^{s}- \\
& -\frac{1}{2}\left(\frac{1}{2 \mu} \sum_{r, s} t_{s}^{r} t_{r}^{s}-\left(\frac{1-2 \mu}{2 \mu}\right) \sum_{r} t_{r}^{r} \sum_{s} t_{s}^{s}\right) \delta_{i}^{m}+F(t) t^{3} . \tag{68}
\end{align*}
$$

We introduce next the "vector" $\eta=\left(\eta_{k}^{i}\right)$ by means of the relation:

$$
\begin{equation*}
G_{i k}=\delta_{k}^{i}+\eta_{k}^{i} . \tag{69}
\end{equation*}
$$

This measures the difference between the pseudo-metric matrix and that corresponding to the identity. Then, we obtain the following estimate for the components of the corresponding inverse matrix $\left(G^{i k}\right):=\left(G_{i k}\right)^{-1}$ :

$$
\begin{equation*}
G^{i k}=\delta_{k}^{i}+\eta_{k}^{i}+F(\eta)\left(\eta^{2}\right) . \tag{70}
\end{equation*}
$$

A straightforward computation shows that the corresponding Christoffel symbols satisfy the following estimate

$$
\begin{equation*}
\Gamma_{k r}^{i}=F(\eta)\left(\eta^{\prime}\right) . \tag{71}
\end{equation*}
$$

Then, it also holds the following estimate

$$
\begin{equation*}
t_{i k}=t_{k}^{i}+F(\eta, t)(\eta t) . \tag{72}
\end{equation*}
$$

For the metric tensor in the deformed "strained" state we have, by definition,

$$
\begin{equation*}
G_{i k}^{*}=G_{i k}+2 \varepsilon_{i k} . \tag{73}
\end{equation*}
$$

Hence, we can also estimate that

$$
\begin{equation*}
G_{i k}^{*}=G_{i k}+\frac{2}{\mu} t_{i k}-2\left(\frac{1-2 \mu}{2 \mu}\right) \sum_{j} t_{j}^{j} \delta_{k}^{i}+F(t, \eta)\left(t^{2}+\eta t\right) . \tag{74}
\end{equation*}
$$

For the tensor with components $c_{k r}^{i}$ measuring the change in the Levi-Civita connections, from the "unstrained" natural state to the deformed "strained" state, we estimate

$$
\begin{equation*}
c_{k r}^{i}=F(\eta, t)\left(t^{\prime}+\eta^{\prime} t\right) . \tag{75}
\end{equation*}
$$

Then we recall that, for the case of Euclidean shells, the above led also to obtain the two following estimates:

$$
\begin{gather*}
\sum_{m} t_{i n ; m}=F(\eta, t)\left(t t^{\prime}+\eta^{\prime} t+\eta t^{\prime}\right)  \tag{76}\\
\sum_{r} t_{h k ; r r}+2 \mu \sum_{r} t_{r r ; h k}=F(\eta, t)\left(\eta t^{\prime \prime}+\left(t^{\prime}\right)^{2}+t \eta^{\prime} t^{\prime}+\eta^{\prime} t^{\prime}+\eta^{\prime \prime} t+\left(\eta^{\prime}\right)^{2} t+\left(\eta^{\prime}\right)^{2} t^{2}+t t^{\prime \prime}\right) \tag{77}
\end{gather*}
$$

We emphasize that the latter expression, obtained in John (1965), was derived on the basis that the ambient of work is precisely Euclidean geometry, where the curvature tensor vanishes. In fact, it was obtained from the corresponding compatibility conditions and, in that context, the curvature is equal to zero for both states of the shell, as expressed in equation (7) of the cited article, i.e.,

$$
\begin{equation*}
0=\mathrm{R}_{a c d b}^{*}=\varepsilon_{a b ; c d}+\varepsilon_{c d ; a b}-\varepsilon_{a d ; c b}-\varepsilon_{b c ; a d}+\sum_{l, s} G_{l s}^{*}\left(c_{a b}^{l} b_{c d}^{s}-c_{a d}^{l} c_{b c}^{s}\right), \tag{78}
\end{equation*}
$$

On the other case, in the present context of affine geometry we have

$$
\begin{equation*}
\mathrm{R}_{a c d b}^{*}=\varepsilon_{a b ; c d}+\varepsilon_{c d ; a b}-\varepsilon_{a d ; c b}-\varepsilon_{b c ; a d}-\frac{1}{2}\left(\sum_{m} G_{a m}^{*} \mathrm{R}_{c b d}^{m}+\sum_{m} G_{c m}^{*} \mathrm{R}_{a d b}^{m}\right)+\sum_{l, s} G_{l s}^{*}\left(c_{a b}^{l} c_{c d}^{s}-c_{a d}^{l} c_{b c}^{s}\right) \tag{79}
\end{equation*}
$$

since, in the present case, the three-dimensional compatibility equations are given in terms of the comparison between the Riemannian curvature tensors of the affine shell, when passing from the natural to the deformed state. Such an equation is obtained by direct application of Lemma 2 in our previous work, see Gigena et al. (2002).

Then, if we denote by $\varepsilon_{a b ; c d}$ the second covariant derivatives with respect to the LeviCivita connection associated to the pseudometric $G$, we further obtain from the latter equation

$$
\begin{equation*}
\varepsilon_{a b ; c d}+\varepsilon_{c d ; a b}-\varepsilon_{a d ; c b}-\varepsilon_{b c ; a d}=-\sum_{l, s} G_{l s}^{*}\left(c_{a b}^{l} b_{c d}^{s}-c_{a d}^{l} c_{b c}^{s}\right)+\mathrm{R}_{a c d b}^{*}-\frac{1}{2}\left(\sum_{m} G_{a m}^{*} \mathrm{R}_{c b d}^{m}+\sum_{m} G_{c m}^{*} \mathrm{R}_{a d b}^{m}\right)( \tag{80}
\end{equation*}
$$

and it is easy to get the following estimates for those tensors

$$
\begin{gather*}
\mathrm{R}=F(\eta)\left(\eta^{\prime \prime}+(\eta)^{2}+\left(\eta^{\prime}\right)^{2}\right)  \tag{81}\\
\mathrm{R}^{*}=F(\eta, t)\left(\left(t^{\prime}\right)^{2}+\left(\eta^{\prime}\right)^{2}+t^{\prime \prime}+\eta^{\prime \prime} t+\eta^{\prime} t^{\prime}+\eta t\right) . \tag{82}
\end{gather*}
$$

In what follows we shall also denote by $\varepsilon$ an upper bound for the absolute values of the principal strains at all points of the shell. Let $P_{0}$ be a point on the undeformed middle surface $M_{0}$ and $D$ the closest affine distance from $P_{0}$ to the lateral surface of the shell. Also, $2 h$ represents the thickness of the shell and $R$ is the typical length associated with the middle surface, all these quantities having been previously introduced above and in our article Gigena et al. (2004).

Then, we introduce the quantity

$$
\begin{equation*}
\theta=\max \left(\frac{h}{D}, \sqrt{\frac{h}{R}}, \sqrt{\varepsilon}\right) \tag{83}
\end{equation*}
$$

and assume, besides, that the circumstances are such that

$$
\begin{equation*}
\theta<\theta_{0}, \tag{84}
\end{equation*}
$$

where $\theta_{0}$ is a constant which depends only on the choice of the strain energy density $W$.
We assume that all of the calculations shall be done for a system of normal affine coordinates, as indicated previously, and also fully described in Gigena et al. (2002). The middle surface of the shell is represented then in the form of Monge's, i.e. as a graph function, where the origin of coordinates is located precisely at $P_{0} \in M_{0}$, with the axis of coordinates chosen to lie at the tangent plane to $M_{0}$ at that point, and with the third axis in the affine normal direction at the same point. The estimates to be computed for the partial derivatives of the function representing $M_{0}$, in the system chosen, shall be used immediately to make the corresponding estimates for the successive covariant derivatives $t_{i k ; r s, . .}$, so that the latter estimates shall be independent of the system originally used.

Thus, we define

$$
\begin{equation*}
\lambda=\frac{\theta_{0} h}{\theta}=\theta_{0} \min \left(D, \sqrt{R h}, \frac{h}{\sqrt{\varepsilon}}\right) \tag{85}
\end{equation*}
$$

and obtain the inequalities

$$
\begin{equation*}
h<\lambda<\sqrt{R h}, h<\frac{R}{4}, \quad \lambda<\frac{D}{2} \tag{86}
\end{equation*}
$$

which are easily seen to be satisfied if we assume, for example, that

$$
\begin{equation*}
\theta_{0}<\frac{1}{2} \tag{87}
\end{equation*}
$$

and from this we obtain that

$$
\begin{equation*}
\frac{1}{R} \leq\left(\theta_{0}\right)^{2} \frac{h}{\lambda^{2}} \leq \theta_{0} \theta \frac{1}{\lambda} \leq \frac{1}{2} \frac{\theta}{\lambda}, \quad \varepsilon<\left(\theta_{0}\right)^{2} \frac{h^{2}}{\lambda^{2}}<\frac{h^{2}}{4 \lambda^{2}} . \tag{88}
\end{equation*}
$$

It is to be further assumed next that $\theta_{0}$ is chosen so small that for the given strain energy function $W$ all of the above formulae are valid in the region defined by

$$
\begin{equation*}
M=\left\{\left(u^{1}, u^{2}, u^{3}\right): \sum\left(u^{\alpha}\right)^{2}<\lambda^{2},\left|u^{3}\right|<h\right\} . \tag{89}
\end{equation*}
$$

Also, from now on we shall use the same symbols of approximation as described in John (1965), represented by $" O$ " and $" ~ O "$, i.e., the first symbol is used in the conventional, classical way except that dependence on $W$ is allowed. Thus, the relation

$$
\begin{equation*}
A=O(B) \tag{90}
\end{equation*}
$$

where $B \geq 0$, means that for a given strain energy function $W$ there exists a positive number
$K$ such that

$$
\begin{equation*}
|A| \leq K B . \tag{91}
\end{equation*}
$$

The second one shall be used in an unconventional sense and only in combination with the first. The relation

$$
\begin{equation*}
A=O(B)+\mathrm{o}(C), \tag{92}
\end{equation*}
$$

where $B \geq 0$ and $C \geq 0$, shall mean that for a given strain energy function $W$ there exists a function $K(k)$, defined for all positive $k$ such that

$$
\begin{equation*}
|A| \leq K(k) B+k C \tag{93}
\end{equation*}
$$

for all $k>0$.
Thus, for example, we may write from previous inequalities that

$$
\begin{gather*}
g_{\alpha \beta, \gamma}=O\left(R^{-1 / 2}\right), \quad g_{\alpha \beta, \eta u}=O\left(R^{-1}\right), g_{\alpha \beta, \mu \nu}=O\left(R^{-3 / 2}\right), \ldots,  \tag{94}\\
B_{\alpha \beta, \gamma}=O\left(R^{-3 / 2}\right), \quad B_{\alpha \beta, \mu}=O\left(R^{-2}\right), B_{\alpha \beta, \mu \nu}=O\left(R^{-5 / 2}\right), \ldots,  \tag{95}\\
\eta=O\left(h^{2} / R^{2}\right), \eta^{\prime}=O\left(1 / R^{5 / 2}\right), \quad \eta^{\prime \prime}=O\left(1 / R^{7 / 2}\right), \ldots, \tag{96}
\end{gather*}
$$

Besides, we shall assume that the strain energy function $W\left(s_{1}, s_{2}, s_{3}\right)$ is defined for all values of $\left|s_{i}\right|$ enough small and is as differentiable as needed. Here $s_{i}$ are the traces of the successive powers of the strain operator. By definition, the "length" of such "strain operator", (1,1)-tensor with components $\varepsilon_{i}^{m}$, is $|\varepsilon|:=\sqrt{\sum_{i, m} \varepsilon_{i}^{m} \varepsilon_{i}^{m}}$. For the metric tensor $G$ sufficiently close to the unit matrix, i.e., for $|\eta|$ sufficiently small, we can estimate $|\varepsilon|$ in terms of the eigenvalues of the matrix $\left(\varepsilon_{i}^{m}\right)$, i.e., in terms of the so-called principal strains.

Then, there exists a positive $\varepsilon_{0}$ only depending on the choice of the strain energy function $W$ such that the strain-stress relations hold for $|\mathcal{E}|<\varepsilon_{0}$, and it also follows that, for such values, $t_{i}^{m}=O(|\varepsilon|)$.

Hence, for a given function $W$ we can also find bounds $t_{0}, \eta_{0}$ such that for $|t|<t_{0}$ and $|\eta|<\eta_{0}$ all of the previously stated estimates are valid and, besides, $|\mathcal{E}|<\varepsilon_{0}$.

## 6 ESTIMATES FOR THE $L_{2}$-NORMS OF SECOND ORDER DERIVATIVES

In what follows, we shall use the following expression of the norm $\|w\|$ for any vector $w=w\left(u^{1}, u^{2}, u^{3}\right)$ defined in the working region $M$ specified above

$$
\begin{equation*}
\|w\|=\sqrt{\iiint_{M}|w| d u^{1} d u^{2} d u^{3}} . \tag{97}
\end{equation*}
$$

The symbol $w^{\prime}$ shall denote the gradient of $w$, i.e., the vector whose components are the first derivatives of the components of $w$ with respect to $u^{1}, u^{2}, u^{3}$. We shall denote, besides, with $w^{\text {" }}$ the "surface" coordinates gradient of $w$, i.e., the vector of first derivatives with respect to $u^{1}, u^{2}$ only. It is well-known that the components of the stress tensor $t_{i k}$ satisfy the symmetry condition $t_{i k}=t_{k i}$. We can represent the estimates obtained from the equations of equilibrium for the Euclidean case, (see Gigena et al. (2003), John (1965) for full details) by

$$
\begin{equation*}
\sum_{m} t_{i n ; m}=P_{i}=F(\eta, t)\left(t t^{\prime}+\eta^{\prime} t+\eta t^{\prime}\right) \tag{98}
\end{equation*}
$$

and the estimates resulting from the compatibility conditions, Gigena et al. (2002), by

$$
\begin{align*}
\sum_{r} t_{h k ; r r}+2 \mu \sum_{r} t_{r r ; h k} & =Q_{h k}  \tag{99}\\
& =F(\eta, t)\left(\eta t^{\prime \prime}+\left(t^{\prime}\right)^{2}+t \eta^{\prime} t^{\prime}+\eta^{\prime} t^{\prime}+\eta^{\prime \prime} t+\left(\eta^{\prime}\right)^{2} t+\left(\eta^{\prime}\right)^{2} t^{2}+t t^{\prime \prime}\right)
\end{align*}
$$

We obtain correspondingly for the Affine Theory of Shells:

$$
\begin{gather*}
\sum_{m} t_{i m ; m}=P_{i}=F(\eta, t)\left(t t^{\prime}+\eta^{\prime} t+\eta t^{\prime}\right),  \tag{100}\\
\sum_{r} t_{h k ; r r}+2 \mu \sum_{r} t_{r r, h k}=Q_{h k}=F(\eta, t)\left(\begin{array}{l}
\left(t^{\prime}\right)^{2}+\eta^{\prime} t t^{\prime}+\left(\eta^{\prime} t\right)^{2}+\left(\eta^{\prime}\right)^{2}+t^{\prime \prime}+\eta^{\prime \prime} t+ \\
+\eta^{\prime} t^{\prime}+\eta t+\eta^{\prime \prime}+(\eta)^{2}+\eta^{\prime \prime} t^{2}+(\eta t)^{2}+ \\
+\left(\eta^{\prime} t\right)^{2}+\eta \eta^{\prime \prime} t+(\eta)^{3} t+\eta t\left(\eta^{\prime}\right)^{2}
\end{array}\right) . \tag{101}
\end{gather*}
$$

In fact, by using the previous estimates one may write:

$$
\begin{equation*}
\varepsilon_{a b ; c d}+\varepsilon_{c d ; a b}-\varepsilon_{a d ; c b}-\varepsilon_{b c ; a d}=\ldots \text { estimated terms, } \tag{102}
\end{equation*}
$$

and, consequently,

$$
\begin{equation*}
t_{h k ; r r}+\frac{1}{1+v} t_{r r ; h k}=\frac{v}{1+v}\left(t_{1, r r} \delta_{h}^{k}-t_{1, r k} \delta_{r}^{h}-t_{1, k r} \delta_{k}^{r}\right)+t_{h r ; k k}+t_{r k ; h r}+. .+ \text { higher order terms }, \tag{103}
\end{equation*}
$$

with

$$
\begin{align*}
t_{i j ; k l}= & t_{i j, k l}-\left(\Gamma_{i k}^{h}\right)_{, l} t_{h j}-\Gamma_{i k}^{h} t_{j, l}-\left(\Gamma_{j k}^{m}\right)_{l} t_{i m}-\Gamma_{j k}^{m} t_{i m, l}-\Gamma_{i l}^{h}\left(t_{h k, j}-\Gamma_{h j}^{r} t_{r k}-\Gamma_{k j}^{s} t_{s h}\right)-  \tag{104}\\
& -\Gamma_{j l}^{m}\left(t_{m i, k}-\Gamma_{m k}^{r} t_{r i}-\Gamma_{i k}^{s} t_{s m}\right)-\Gamma_{k l}^{q}\left(t_{q i, j}-\Gamma_{q i}^{r} t_{r j}-\Gamma_{i j}^{s} t_{s q}\right)
\end{align*} .
$$

In the two latter equations, and also in what follows, we have used the summation convention whenever repeated indices appear.

Finally, by using all of the above expressions one may write, for the case of Affine Shells, estimates which resemble the ones obtained for the Euclidean case.

In fact, we develop estimates for the norms of the second derivatives of the (symmetric) tensor components $t_{i k}$, in terms of the above quantities $P_{i}$ and $Q_{h k}$, introducing the auxiliary
function $\phi: M \rightarrow \mathbb{R}$, defined by:

$$
\begin{equation*}
\phi=\left(1-\lambda^{-2} u^{\alpha} u^{\alpha}\right)^{2} . \tag{105}
\end{equation*}
$$

One verifies easily the following properties

$$
\begin{gather*}
0 \leq \phi \leq 1 \quad \text { and } \quad \partial_{3} \phi=0  \tag{106}\\
\left|\phi^{\bullet}\right|=\sqrt{\partial_{\alpha} \phi \partial_{\alpha} \phi} \leq \frac{4}{\lambda} \phi^{1 / 2}  \tag{107}\\
\left|\phi^{\cdot \bullet}\right|=\sqrt{\partial_{\alpha \beta} \phi \partial_{\alpha \beta} \phi} \leq \frac{4 \sqrt{10}}{\lambda^{2}} . \tag{108}
\end{gather*}
$$

Next, we consider the integral

$$
\begin{equation*}
A=: \sqrt{\iiint_{M} \phi^{2} \partial_{i \alpha} t_{3 k} \partial_{i \alpha} t_{3 k} d u^{1} d u^{2} d u^{3}} . \tag{109}
\end{equation*}
$$

Whereas, in a similar fashion to the Euclidean case one obtains the estimates

$$
\begin{equation*}
\|\phi t \cdot \bullet\|=O\left(A+\left\|\phi P^{\prime}\right\|+\|\phi Q\|+\lambda^{-2}\|t\|\right) \tag{110}
\end{equation*}
$$

with

$$
\begin{equation*}
t^{\bullet \bullet}=\left(\partial_{\alpha \beta} t_{i k}\right) ; \quad P^{\prime}=\left(\partial_{k} P_{i}\right) ; \quad Q=\left(Q_{i k}\right) \tag{111}
\end{equation*}
$$

and

$$
\begin{equation*}
\left\|\phi t^{\prime \prime}\right\|=O\left(\lambda h^{-1}\left(A+\left\|\phi P^{\prime}\right\|+\|\phi Q\|+\lambda^{-2}\|t\|\right)\right) . \tag{112}
\end{equation*}
$$

Similarly, one may estimate the above expression $A$ :

$$
\begin{equation*}
A=O\left(\lambda^{-2}\|t\|+\|\phi Q\|+\lambda h^{-1}\left\|\phi P^{\bullet}\right\|\right)+o\left(\left\|\phi t^{\bullet \bullet}\right\|+\lambda^{-1} h\left\|\phi t^{\prime \prime}\right\|\right) \tag{113}
\end{equation*}
$$

Then, it follows that

$$
\begin{gather*}
\left\|\phi t^{\bullet}\right\|=O\left(\lambda^{-2}\|t\|+\|\phi Q\|+\left\|\phi^{\bullet} t^{\bullet}\right\|^{2}+\lambda h^{-1}\left\|\phi P^{\bullet}\right\|\right)  \tag{114}\\
\left\|\phi t^{\prime \prime}\right\|=O\left(\lambda h^{-1}\left(\lambda^{-2}\|t\|+\|\phi Q\|+\left\|\phi^{\bullet} t^{\bullet}\right\|^{2}+\lambda h^{-1}\left\|\phi P^{\bullet}\right\|\right)\right) . \tag{115}
\end{gather*}
$$

Next, we make use of the particular form of the second and third members in equations (100) and (101):

$$
P_{i}=F(\eta, t)\left(t t^{\prime}+\eta^{\prime} t+\eta t^{\prime}\right), Q_{h k}=F(\eta, t)\left(\begin{array}{l}
\left(t^{\prime}\right)^{2}+\eta^{\prime} t t^{\prime}+\left(\eta^{\prime} t\right)^{2}+\left(\eta^{\prime}\right)^{2}+t^{\prime \prime}+\eta^{\prime \prime} t+  \tag{116}\\
+\eta^{\prime} t^{\prime}+\eta t+\eta^{\prime \prime}+(\eta)^{2}+\eta^{\prime \prime} t^{2}+(\eta t)^{2}+ \\
+\left(\eta^{\prime} t\right)^{2}+\eta \eta^{\prime \prime} t+(\eta)^{3} t+\eta t\left(\eta^{\prime}\right)^{2}
\end{array}\right),
$$

as well as on the respective (gradient) derivatives $P^{\prime}, P^{\bullet}$.

Thus, by having into account these differences with the Euclidean case we find, by taking at the same time a sufficiently small value for the constant $\theta_{0}$, that

$$
\begin{align*}
& \|\phi t \cdot\|=O\left(\lambda^{-2}\|t\|^{2}\right)  \tag{117}\\
& \left\|\phi t^{\prime \prime}\right\|=O\left(\lambda^{-1} h^{-1}\|t\|^{2}\right) \tag{118}
\end{align*}
$$

## 7 ESTIMATES FOR THE $L_{2}$-NORMS OF HIGHER ORDER DERIVATIVES

From the previous development we observe that a similar approach can be made by considering again the differential equations (100) and (101), by taking partial derivatives, i.e., for a fix value of $\gamma$, the $\partial_{\gamma} t_{i k}$ quantities satisfy the same equations if we substitute $P_{i}$ y $Q_{b k}$ respectively by $\partial_{\gamma} P_{i}$ and $\partial_{\gamma} Q_{h k}$. Thus, we obtain estimates for $\left\|\phi \partial_{\gamma \alpha \beta} t_{i k}\right\|$ and $\left\|\phi \partial_{\gamma r s} t_{i k}\right\|$. On the other hand, the remaining component $\left\|\phi \partial_{333} t_{i k}\right\|$ may also be estimated if one observes that all those of the form $\partial_{333} t_{i k}$ are expressible in terms of the quantities $\partial_{\gamma r s} t_{m n}$ and suitable derivatives of the expressions $P_{i}$ and $Q_{h k}$. In such a fashion we are led to obtain estimates like the following ones:

$$
\begin{gather*}
\left\|\phi t^{\cdots \cdots}\right\|=O\left(\lambda^{-2}\left\|t^{\bullet}\right\|+\left\|\phi Q^{\prime}\right\|+\left\|\phi P^{\prime \prime}\right\|+\lambda h^{-1}\left\|\phi P^{\bullet \bullet}\right\|\right)  \tag{119}\\
\left\|\phi t^{\prime \prime \prime}\right\|=O\left(\lambda h^{-1}\left(\lambda^{-2}\left\|t^{\bullet}\right\|+\left\|\phi Q^{\prime}\right\|+\left\|\phi P^{\prime \prime}\right\|+\lambda h^{-1}\left\|\phi P^{\bullet \bullet}\right\|\right)\right) \tag{120}
\end{gather*}
$$

So far, in all of the previous development we kept fixed the value of $\lambda$ regarding the defining region $M$, the function $\phi$, the norm $\|\omega\|$, all of those depending on $\lambda$. However, for estimating higher order estimates we shall need to reduce further and further the region of work. Thus, we are led to rename all of those objects, for example by calling $M_{\lambda}$ instead of $M, \phi_{\lambda}$ in place of $\phi$, and $\|\omega\|_{\lambda}$ for $\|\omega\|$. Moreover, we shall also proceed in our work by replacing $\lambda$ for $\lambda / 2$. Consequently, we replace too $\theta_{0}$ for $\theta_{0} / 2$, i.e., in doing so we restrict now to values $\theta$ such that $\theta<\theta_{0} / 2$. Next, we observe that $1=O\left(\phi_{\lambda}\right)$ in the region $M_{1 / 2}$, so that:

$$
\begin{equation*}
\|\omega\|_{\lambda / 2}=O\left(\left\|\phi_{\lambda} \omega\right\|_{\lambda}\right) \tag{121}
\end{equation*}
$$

It follows that

$$
\begin{equation*}
\left\|\phi_{1 / 2} t \cdots\right\|_{2 / 2}=O\left(\lambda^{-2}\left\|\phi_{\lambda} t^{\bullet}\right\|_{\lambda}+\left\|\phi_{1 / 2} Q^{\prime}\right\|_{2 / 2}+\left\|\phi_{1 / 2} P^{\prime \prime}\right\|_{2 / 2}+\lambda h^{-1}\left\|\phi_{1 / 2} P^{\cdot}\right\|_{\lambda / 2}\right) \tag{122}
\end{equation*}
$$

with the corresponding expression for $\left\|\phi_{1 / 2} t^{\prime \prime \prime}\right\|_{2 / 2}$ and also the corresponding changes for the (gradient) derivatives $P^{\prime \prime}$ and $P^{\bullet \bullet}$.

Then, a similar argument to the one in the Euclidean case allows us to conclude first that:

$$
\begin{align*}
& \left\|\phi_{1 / 2} t^{\prime \prime}\right\|_{1 / 2}=O\left(\lambda^{-2} h^{-1}\|t\|_{\lambda}^{2}\right),  \tag{123}\\
& \left\|\phi_{1 / 2} \cdots\right\|_{\lambda / 2}=O\left(\lambda^{-3}\|t\|_{\lambda}^{2}\right) . \tag{124}
\end{align*}
$$

This kind of process can be repeated over and over again, with the corresponding changes. For example, by applying the relations (100) and (101) to $t^{\bullet \bullet}$, in place of $t$ and, simultaneously replacing $\lambda$ by $\lambda / 4, P$ by $P^{\bullet \bullet}$ and $Q$ by $Q^{\bullet \bullet}$, we find that

$$
\begin{equation*}
\left\|\phi_{1 / 4} t^{\cdots \cdots}\right\|_{2 / 4}=O\left(\lambda^{-2}\left\|\phi_{1 / 2} t^{\bullet}\right\|_{2 / 2}+\left\|\phi_{1 / 4} Q^{\prime \prime}\right\|_{2 / 4}+\left\|\phi_{1 / 4} P^{\prime \prime \prime}\right\|_{2 / 4}+\lambda h^{-1}\left\|\phi_{1 / 4} P^{\cdots}\right\|_{1 / 4}\right), \tag{125}
\end{equation*}
$$

with a similar expression for $\left\|\phi_{\lambda_{1}} t^{\prime \prime \prime}\right\|_{\lambda / 4}$.
Therefore, we obtain the following estimates for partial derivatives of higher order

$$
\begin{gather*}
\partial_{k_{1}, k_{2}, i_{1}, i_{2}, \ldots, i_{n}} t=O\left(\varepsilon^{2} \lambda^{1-n} h^{-1}\right),  \tag{126}\\
\partial_{k_{1}, k_{2}, \alpha_{1}, \alpha_{2}, \ldots, \alpha_{n}} t=O\left(\varepsilon^{2} \lambda^{-n}\right) . \tag{127}
\end{gather*}
$$

On the other hand by using the estimate expressed in equation (71), i.e., $\Gamma_{k r}^{i}=F(\eta)\left(\eta^{\prime}\right)$, and its successive derivatives, together with the ones previously determined in (96), i.e.,

$$
\eta=O\left(h^{2} / R^{2}\right), \eta^{\prime}=O\left(1 / R^{5 / 2}\right), \eta^{\prime \prime}=O\left(1 / R^{7 / 2}\right), \ldots
$$

we come to the conclusion that the same kind of estimates are valid for the corresponding covariant derivatives respect to the Levi-Civita connection:

$$
t_{k_{1}, k_{2} ; i, i, i_{2}, \ldots, i_{n}}=O\left(\varepsilon^{2} \lambda^{1-n} h^{-1}\right)
$$

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